

Image Processing and Imaging

Image Restoration

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1 Introduction

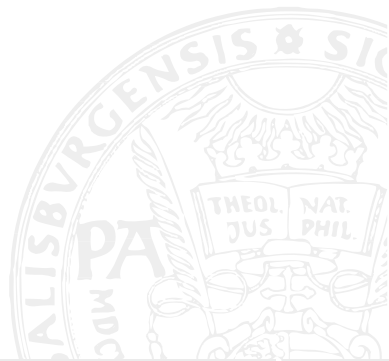
2 Image Distortion

3 Distortion Determination

- Image Analysis
- Experimental Distortion Determination
- Distortion Determination by Modeling

4 Distortion Removal

5 Wiener Filtering



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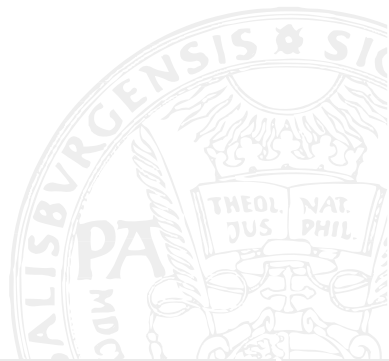


Image Restoration Introduction

- Improves image quality in case of an existing image distortion which should be removed
- Contrasting to image enhancement the aim is to restore a (virtual) original image
- Either the type of distortion is known or it has to be estimated
- Reasons for existing distortions:
 - Defocus, motion blur, noise (transmission errors, sensor noise, ...), defects in the optical system (Hubble), etc.

Deterministic methods: for images with low amount of noise and known distortion function

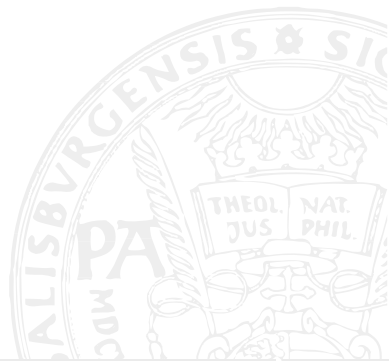
Stochastic methods: try to identify the best restoration with respect to some statistical error-criterion (e.g. least-squares criterion)

The better the distortion is known, the better it can be removed and the better is the resulting restoration. In many cases, the distortion needs to be estimated:

A priori estimation: distortion is known or is obtained before the restoration starts

A posteriori estimation: image analysis based on *interesting* pixels (e.g. edges, straight lines, homogeneous areas) and it is attempted to estimate / reconstruct their original properties.

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We assume a position invariant linear distortion and independent additive noise:

$$g(x, y) = h(x, y) * f(x, y) + v(x, y) \quad (1)$$

$v(x, y)$... noise

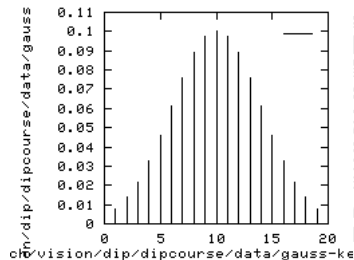
h ... distortion (position invariant)

Distorted signal can be represented in the DFT domain as follows (linearity and convolution theorem):

$$\hat{g}(u, v) = \hat{h}(u, v) \cdot \hat{f}(u, v) + \hat{v}(u, v)$$



smoothed image



Gauss kernel (1-D)

Figure: Example: Smoothing as image distortion

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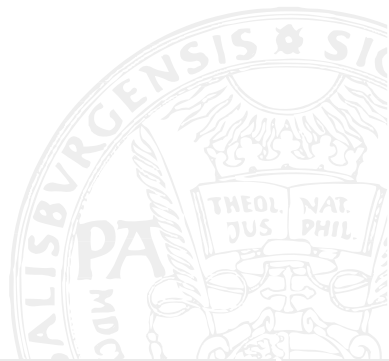
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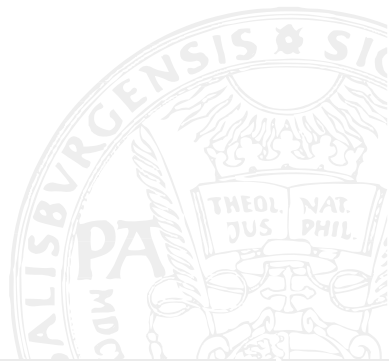
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- Several alternatives how to determine (estimate, approximate) the image distortion present
- All the following techniques are approximative by nature → ideas are called “blind deconvolution”

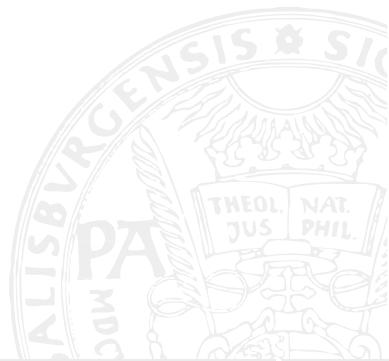
Image Analysis

- In the distorted image we choose image regions with “obvious” image content, e.g. a sharp edge
- Corresponding image part is denoted $g_s(x, y)$
- Create an approximation $f_s^a(x, y)$ for this image part of the original image
- Selection of the region → noise should not affect the result too much
- We are able to compute the distortion function in the DFT domain for the region selected:

$$\hat{h}_s(u, v) = \frac{\hat{g}_s(u, v)}{\hat{f}_s^a(u, v)}$$

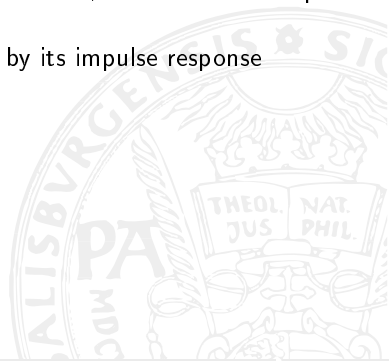
- Due to the assumed position invariance, the distortion function can be generalised to the entire image

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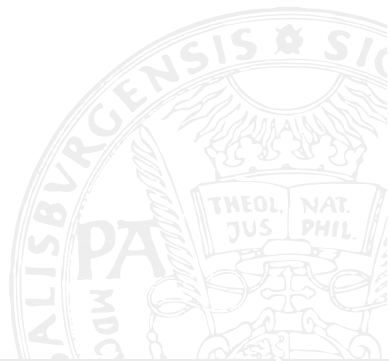


- Assume that the equipment used to take the distorted image is available (or at least the same or similar model)
- Capture an image similar to the one subject to restoration
- Try to generate a distortion highly similar to the one to model by systematic testing of the system configurations (e.g. different camera settings)
- Once identified, we take an image of an intensive point light source, to obtain the impulse response of the distortion
 - A distortion of the considered type is uniquely characterised by its impulse response
- DFT of an impulse is a constant A , thus we get:

$$\hat{h}(u, v) = \frac{\hat{g}(u, v)}{A}$$



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Here we exploit knowledge about models of physical processes

Examples for simple distortions are:

Relative (uniform) motion between camera and object:

- $f(x, y)$ moves in a way such that $x_0(t)$ and $y_0(t)$ are the time-dependent motion components in x and y direction
- Entire image exposition is obtained by integrating the image function over the time-frame of the shutter being opened T :

$$g(x, y) = \int_0^T f(x - x_0(t), y - y_0(t)) dt$$

$$\hat{g}(u, v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(x, y) e^{-2\pi i(ux+vy)} dx dy$$

$$\hat{g}(u, v) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left[\int_0^T f(x - x_0(t), y - y_0(t)) dt \right] e^{-2\pi i(ux+vy)} dx dy$$

- Integration order can be swapped:

$$\hat{g}(u, v) = \int_0^T \left[\int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(x - x_0(t), y - y_0(t)) e^{-2\pi i(ux + vy)} dx dy \right] dt$$

- Expression between $[]$ is the DFT of the shifted function $f(x - x_0(t), y - y_0(t))$
- Using the translation properties of the DFT and the independence between $\hat{f}(u, v)$ and t we obtain:

$$\hat{g}(u, v) = \int_0^T \hat{f}(u, v) e^{-2\pi i(ux_0(t) + vy_0(t))} dt = \hat{f}(u, v) \int_0^T e^{-2\pi i(ux_0(t) + vy_0(t))} dt$$

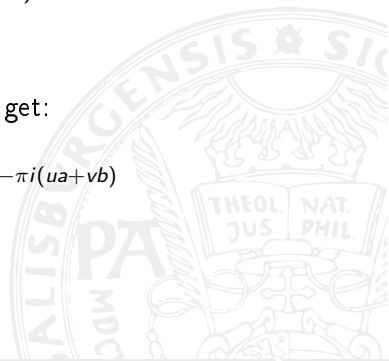
- Thus, we set $\hat{h}(u, v) = \int_0^T e^{-2\pi i(ux_0(t) + vy_0(t))} dt$

- Setting for example $x_0(t) = at/T$ and $y_0(t) = 0$ we get motion only in x-direction (e.g. taking pictures from a moving car)
- At time $t = T$ the image moved by distance a . We get

$$\hat{h}(u, v) = \int_0^T e^{-2\pi i u a t/T} dt = \frac{T}{\pi u a} \sin(\pi u a) e^{-\pi i u a}$$

- For two-dimensional motion (also $y_0(t) = bt/T$ as well) we get:

$$\hat{h}(u, v) = \frac{T}{\pi(ua + vb)} \sin(\pi(ua + vb)) e^{-\pi i(ua + vb)}$$



Defocus:

$$\hat{h}(u, v) = \frac{J_1(ar)}{ar} \quad \text{mit } r^2 = u^2 + v^2$$
$$J_1(x) = \sum_{k=0}^{\infty} \frac{(-1)^k (x/2)^{2k+1}}{k!(k+1)!}$$

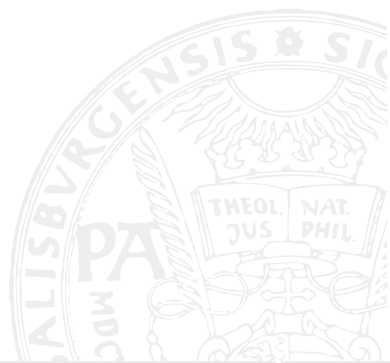
J_1 ... Bessel function Order 1

a ... Extent of Defocus

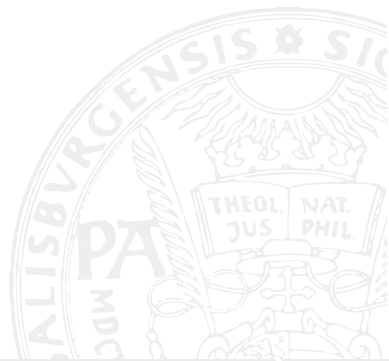
Atmospheric Turbulence:

$$\hat{h}(u, v) = e^{-c(u^2+v^2)^{5/6}}$$

c is determined experimentally



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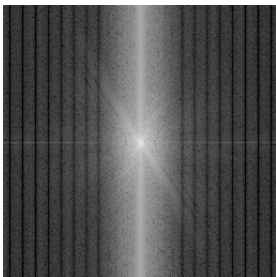
Inverse Filtering:

- Need to construct a restoration filter, exhibiting a transfer function invers to the distortion: $\hat{h}^{-1}(u, v)$
- The resulting procedure is denoted “Inverse Filtering”:

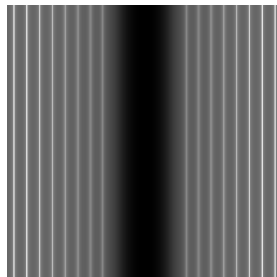
$$\hat{f}(u, v) = \hat{g}(u, v) \cdot \hat{h}^{-1}(u, v) - \hat{v}(u, v) \cdot \hat{h}^{-1}(u, v)$$

- If the noise contribution is not too high, restoration is identical to inverse convolution
- If the noise contribution is too high or $\hat{h}(u, v)$ is too small we result in a large value for $\hat{v}(u, v) \cdot \hat{h}^{-1}(u, v)$
 - Dominates the inverse filtering
- Result of inverse filtering with high noise contribution (see next slide)
- Smoothed image (originally given as float data type) has been casted to char data type, resulting in significant (quantisation) noise
- Result of inverse filtering is quite poor!

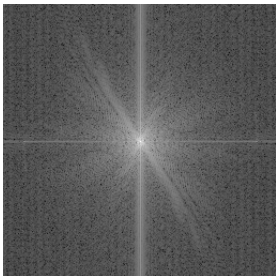
Distortion Removal - Inverse Filtering Example



DFT of the smoothed image



DFT of the Gauss kernels



DFT after inverse filtering

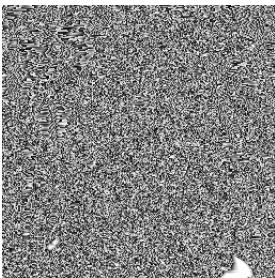


image after restauration

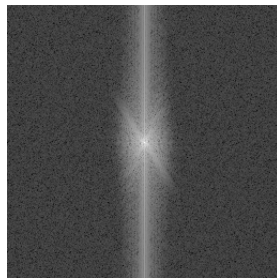
Distortion Removal - Inverse Filtering High Noise Example



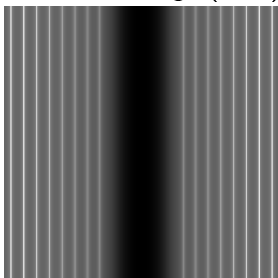
smoothed image (char)



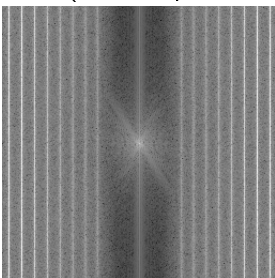
difference (with respect to float)



DFT of smoothed img (Byte)



DFT of Gauss kernel



DFT after inverse filtering

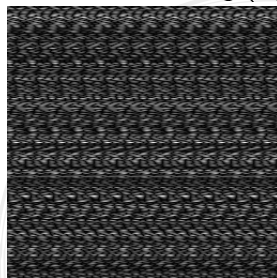


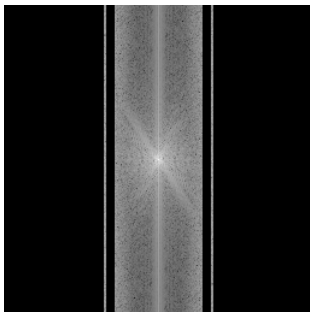
image after restoration

Pseudo-Inverse Filtering:

$$\hat{h}^{-1}(u, v) = \begin{cases} h^{-1}(u, v) & \text{if } |\hat{h}(u, v)| > T \\ 0 & \text{if } |\hat{h}(u, v)| \leq T \end{cases}$$

Case of $\hat{h}(u, v)$ being too small is “corrected” to prevent a large value for $\hat{v}(u, v) \cdot \hat{h}^{-1}(u, v)$

Still, large values for $\hat{v}(u, v)$ remain unsolved (noise contribution)



DFT after pseudo-inverse Filtering



Image after pseudo-inverse Filtering

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Wiener Filtering:

- Additionally exploits a priori knowledge about the noise contribution
- Delivers an estimation of the non-distorted image \hat{f} with minimal error $f(i,j) - \hat{f}(i,j)$ with respect to some determined metric

$$\hat{f}(u, v) = \hat{h}_w(u, v) \cdot \hat{g}(u, v) \quad (2)$$

$$\hat{h}_w(u, v) = \frac{\hat{h}^*(u, v)}{|\hat{h}(u, v)|^2 + \frac{s_{xx}(u, v)}{s_{\eta\eta}(u, v)}} \quad (3)$$

- s_{xx} and $s_{\eta\eta}$ are the spectral densities of the noise and the original (non-distorted) image
 - Of course, these values are difficult to obtain
- To be able to conduct this type of filtering, information about the distortion and statistical knowledge about the noise are required

- Actual knowledge required for this process (e.g. computation of the spectral density of the noise) might be hard or impossible to obtain
- A “parameterised” Wiener filter can be employed:

$$\hat{h}_w^K(u, v) = \frac{\hat{h}^*(u, v)}{|\hat{h}(u, v)|^2 + K}$$

- In the process, K is optimised until the best result is reached
- Can be improved by using so-called “constrained least square” filters
- Intuitively this means that the optimisation of K is done with respect to some least squares criterion
 - E.g. to maximise image smoothness (or minimisation of image variations), expressed in terms of a gradient operator

Original



distorted image



Wiener filtering



Figure: Wiener Filtering

